

# Clusters under strong vuv pulses: A quantum-classical hybrid description incorporating plasma effects

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The quantum-classical hybrid description of rare-gas clusters interacting with intense light pulses which we have developed is described in detail. Much emphasis is put on the treatment of screening electrons in the cluster which set the time scale for the evolution of the system and form the link between electrons strongly bound to ions and quasifree plasma electrons in the cluster. As an example, we discuss the dynamics of an Ar<sub>147</sub> cluster exposed to a short vuv laser pulse of 20 eV photon energy.

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## I. INTRODUCTION

The interaction of rare-gas clusters with intense laser pulses has attracted considerable attention over the last decade since spectacular experiments showed how effective energy from the laser beam can be coupled into the cluster, culminating in the demonstration of fusion in exploding deuterium clusters [1]. Other efforts aim at understanding the interaction of matter beyond a few atoms with light of different wavelengths, motivated by the fact that little is known about such processes with intense light of frequencies different than the infrared range around 800 nm wavelength. First experiments at FLASH (Free-electron LASer in Hamburg) have revealed very efficient energy absorption at 98 nm wavelength [2] and have triggered theoretical research [3–6] as recently reviewed [7].

Clearly, the time-dependent description of a finite many-body system interacting with a laser pulse is challenging. A full quantum description is numerically impossible and would probably provide little insight conceptually. One can observe two different strategies in the literature to tackle this problem: (i) A more or less full quantum approach for a single atom in the cluster treating the influence of the other atoms and ions in the cluster more approximately as a kind of environment [3,8]. (ii) A classical propagation of all charged particles in time (ions and electrons) with quantum mechanical elements added in the form of rates at various levels of sophistication [9–16].

We have followed the latter route, describing the motion of all charged particles classically, while they interact and are subjected to the external dipole-coupled laser field. Initial bound electronic motion is not treated explicitly, but only in the form of (quantum) ionization rates which are integrated into the classical dynamics with a Monte Carlo scheme [7]. The approach works very well, as long as the electrons, once ionized from their mother ions, behave essentially classically. This is the case if they are subject to a strong external field (as is in the case with strong ir pulses) or are instantaneously free (as in the case of hard x rays). Under vuv radiation, however, the photoelectrons stay often very close to their mother-ions or other ions, effectively screening them and modifying subsequent ionization processes.

The reason for this behavior is the size of the quiver motion (typically less than a ground-state electron orbit of the

rare-gas atom) and the small kinetic energy which remains of the photon energy in excess of the ionization potential. The latter is not the case for hard x rays where the photon energy is high enough to remove the photoelectrons completely from the cluster (at least for moderate cluster sizes). For intense ir fields, on the other hand, the photon energy is much too low for the photoelectrons to leave the cluster instantaneously, but the quiver motion is (for typical intensities) of the size of the cluster or even larger implying that the photoelectrons certainly will not remain in the vicinity of specific ions. Rather, they are dragged back and forth through the cluster by the laser field. Hence, the photoionization of ions surrounded by screening electrons is a phenomenon unique to vuv radiation, adding to the challenge of a theoretical description. On the other hand, as will become clear subsequently, exactly those electrons which screen individual ions define a time scale, suitable to formulate a coarse-grained electron dynamics in the cluster. It is the key to incorporate physical processes which in our approach lie at the interface of the classical and quantum descriptions, such as the influence of the surrounding charged particles (ions and electrons) on the photoionization rate of an ion in the cluster.

In the next section we summarize our quantum-classical hybrid description which contains quantum photoionization rates of multielectron ions, classical propagation, and, as an element in between, the treatment and identification of those electrons which screen the ions. In Sec. III we explain how to deal with photoionization of bound electrons into the plasma of cluster electrons. Section IV discusses as an application and illustrative example the illumination of Ar<sub>147</sub> with a vuv pulse of 62 nm wavelength. The paper ends with a short summary in Sec. V.

## II. HYBRID DESCRIPTION OF TIME-DEPENDENT CLUSTER EVOLUTION

The interaction of a cluster with intense radiation can be partitioned into three parts: (A) atomic ionization, (B) cooperative interaction of ions among each other and with electrons, and (C) relaxation. During phase (A) which lasts approximately until every second atom in the cluster is (singly) ionized, one can simply apply the atomic photoionization cross sections for many-electron atoms supplied, e.g., in

[17,18]. Gradually, the bound electrons feel the cluster environment, which has roughly three effects: (i) the ionization potential is lowered through close by ions, (ii) previously ionized electrons trapped by the cluster screen the ion under consideration, and (iii) the global field, generated by electrons and ions, modifies the ionization potential as well. We will treat all three effects which happen during phase (B) of the dynamics in Sec. III. Here, we describe briefly how to extrapolate the known photoionization cross sections below the ionization threshold.

### A. Atomic photoionization

The calculation of photoionization and excitation for individual energy levels of isolated atoms and ions is straightforward. Including (quantum) rates into the classical propagation of charged particles in the cluster, we do not resolve the angular dependences of the photoionization or photoexcitation cross section. Averaging over angular degrees of freedom considerably simplifies the cross sections which only depend on the radial wave functions for the respective mean initial and final energies of the phototransition; for a similar philosophy, see [6]. We start from the dipole matrix element for linearly polarized light along the  $\hat{z}$  direction and consider the transition between initial and final atomic states with well-defined orbital quantum numbers  $lm$ :

$$\langle f|z|i\rangle = d_r(\ell_f, \ell_i) \int d\Omega Y_{\ell_f m_f}^*(\Omega) \cos \theta Y_{\ell_i m_i}(\Omega), \quad (1)$$

with the radial dipole matrix element

$$d_r(\ell_i, \ell_f) = \int_0^\infty dr u_{\ell_f}(r) r u_{\ell_i}(r). \quad (2)$$

Within the independent-particle picture all states in a shell defined by orbital angular momentum  $\ell_i$  are degenerate and have the same radial wave function  $u_{\ell_i}(r)$ . It is an eigenfunction to a spherically symmetric effective single-particle potential obtained (along with the eigenfunction) from the Cowan code [17]. Within this mean-field approximation the full photoabsorption cross section is easily obtained by summing over all available final states and averaging over the initial states in the shell  $\ell$ :

$$\sigma_\ell(\omega) = \frac{w_i}{3} \frac{4\pi^2 \alpha \omega}{2\ell + 1} [(\ell + 1)d_r^2(\ell, \ell + 1) + \ell d_r^2(\ell, \ell - 1)]. \quad (3)$$

Here,  $w_i$  gives the number of available initial electrons.

Furthermore, in the cluster the ions are surrounded by other electrons and ions. The latter lower the potential barriers for ionization into the cluster. Hence, what is discrete photoexcitation in an isolated atom becomes at the same photon energy ionization into the cluster. Therefore, we need an interpolation of the discrete excitation spectrum distributing the oscillator strength over the respective photon energy interval.

In the following, we compare two different interpolations. In the first one [19], approximate analytical expressions for

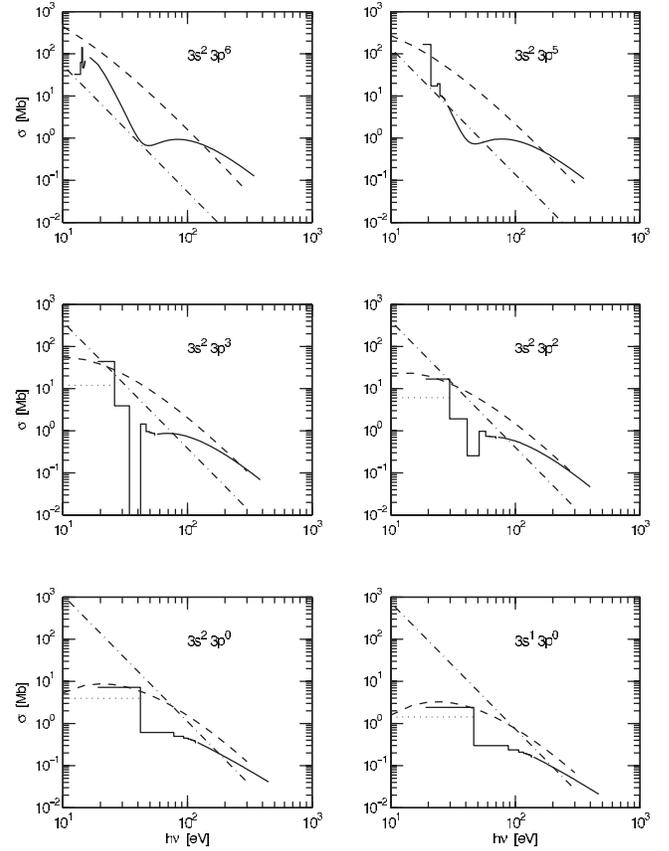


FIG. 1. Atomic photoionization cross sections extrapolated below the threshold according to [19] (dashed lines) and by interpolation of the discrete spectrum (solid lines). The dash-dotted lines are an approximation with hydrogenic wave functions; see text. The initial configuration is displayed in each panel.

the photoabsorption of many-electron atoms are derived. The corresponding results for the Ar atom are shown in Fig. 1 with dashed lines. For comparison, the cross sections with hydrogenic wave functions [20] are also shown. In the second approximation we define a continuous photoabsorption in the region of discrete spectral lines by demanding that the renormalized photoexcitation cross section merge smoothly with the photoionization cross section at threshold [21]. This is achieved by distributing the oscillator strength  $f_n$  of a spectral line  $E_n$  over an interval halfway to each of the adjacent lines such that

$$\sigma_n(\omega) = \frac{2\pi\alpha f_n}{(E_{n+1} - E_{n-1})/2}, \quad (4)$$

where  $\sigma_n(\omega)$  is now the interpolated photoabsorption cross section, Eq. (3), for  $(E_{n-1} + E_n)/2 < \omega < (E_n + E_{n+1})/2$ . The result (solid line in Fig. 1) shows reasonable agreement with the analytical approximation for low and high energies. For intermediate energies, the well-known Cooper minima lead to considerably lower values than in the analytical approximation which does not account for this interference effect. Hence, we will use in the following the approximation (4).

There is a special case for Eq. (4)—namely, ionization just above the threshold. In this case, the oscillator strength of the first spectral line  $n_0$  above the threshold has to be distributed on the whole interval between the ionization threshold,  $\Delta E_{\text{ip}}^{\text{eff}}$ , and the next higher line  $n_0+1$ :

$$\sigma_{n_0}(\omega) = \frac{2\pi\alpha f_{n_0}}{(E_{n_0+1} + E_{n_0})/2 - \Delta E_{\text{ip}}^{\text{eff}}}. \quad (5)$$

As will be shown later in Sec. III C,  $\Delta E_{\text{ip}}^{\text{eff}}$  depends strongly on the cluster environment. The dotted line in Fig. 1 depicts the lower bound for the photoionization cross section which corresponds to the extreme case when the ionization potential has been lowered to its smallest value.

### B. Classical propagation for the Coulomb interaction under vuv radiation

The propagation of classical particles is in principle straightforward. More refined methods, such as tree codes are only worth the effort of coding for large clusters ( $10^5$  electrons and ions and more). Another issue is the Coulomb interaction. Using the real Coulomb potential with its singularity is numerically very costly (small time steps close to the singularity) and leads for more than two bound electrons to artificial autoionization since one electron can fall into the nucleus (below the quantum ground-state energy) and another one can be ionized with the released energy. In strong-field physics (at ir frequencies) the so-called soft-core potential

$$U(r) = -\frac{Z}{(r^2 + a^2)^{1/2}} \quad (6)$$

has been used routinely where the singularity is cut off by the smoothing parameter  $a$  chosen to get potential depths (slightly) below the true ionization potential of the atom or ion. As long as the quiver amplitude  $x_{\text{quiv}} = F/\omega^2 \gg a$ , the cutoff is irrelevant for the dynamics, as is typically the case for a strong pulse (for  $3 \times 10^{16}$  W/cm<sup>2</sup> and 800 nm wavelength,  $x_{\text{quiv}} \approx 500a_0$ , while  $a$  is of the order of  $1a_0$ ). However, at 10 times higher photon frequency and a factor of 100 weaker peak intensity—which is realistic, e.g., for the FLASH source— $x_{\text{quiv}} \approx a$ . Even more problematic is the fact that the soft-core potential is harmonic about its minimum with a characteristic frequency  $(Z/a^3)^{1/2}$  which could become resonant with the vuv laser frequency.

To avoid these problems we use a different approximative potential which has the correct asymptotic Coulomb behavior at large distances, but lacks an eigenfrequency since it has a nonzero slope at  $r=0$ :

$$V(r) = -\frac{Z}{r}(1 - e^{-r/a}). \quad (7)$$

Here,  $a$  is chosen in analogy to the potential (6) discussed above. Note that  $U(r \rightarrow 0) = V(r \rightarrow 0) = -Z/a$ . For rare-gas atoms  $a$  is of the order of 1:  $a=1.74$  for Xe and  $a=1.4$  for Ar.

For the photon frequency used here ( $\hbar\omega=20$  eV) there is no qualitative difference using a U-shape (soft-core) or a

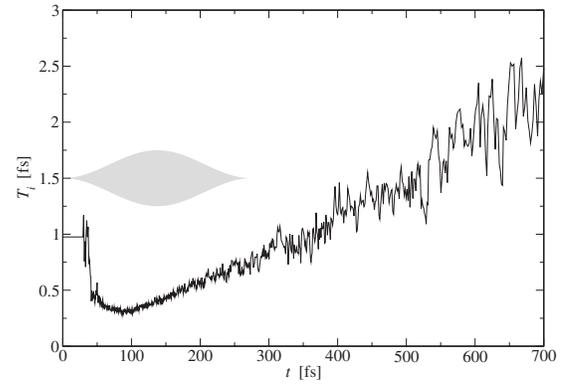


FIG. 2. Time intervals  $T_i$  of the averaged localized motion of the screening electrons as a function of real time for an  $\text{Ar}_{147}$  cluster with laser parameter as specified in Sec. IV.

V-shape potential. The subsequent considerations also do not depend on the approximate or exact form of the of the Coulomb potential. Therefore we will use the generic form  $v(i, j)$ , which could be either of the two options

$$v(i, j) = \begin{cases} [(\mathbf{r}_i - \mathbf{r}_j)^2 + a^2]^{-1/2} & \text{U shape,} \\ (1 - e^{-|\mathbf{r}_i - \mathbf{r}_j|/a})|\mathbf{r}_i - \mathbf{r}_j|^{-1} & \text{V shape.} \end{cases} \quad (8)$$

Our numerical examples presented here have been obtained with the V-shape potential.

### C. Identification of localized and delocalized electrons

At photon energies comparable to or less than the ionization potential of a cluster ion ( $\hbar\omega \leq 30$  eV), most of the photoelectrons remain in the cluster; i.e., quasifree electrons are produced. They thermalize quickly; i.e., they form a plasma.

We have to determine which among these quasifree electrons travel all over the cluster and visit many ions (delocalized quasifree electrons) and which revolve about a single ion—that is, are effectively in excited states about an ion (localized quasifree electrons). To do so, we record the revolution angle  $\phi(t)$  of each classical electron about its closest ion as a function of time. If the electron  $j$  moves for two revolutions ( $\phi_j = 4\pi$ ) about the same ion  $\alpha$ , we consider it as localized, and the period of its motion  $T^{j\alpha}$  is then given by  $2\pi = \phi_{j\alpha}(T^{j\alpha})$ .

The average period  $\bar{T}^{j\alpha}$  of all localized classical electrons sets the time scale for a coarse-grained dynamics. This time scale changes slowly in real time  $t$  due to changes in number and energy of the localized electrons; see Fig. 2. With an initial guess for the first averaged period  $T_0=1$  fs, we update  $T$  after a time  $t=T_i$  according to the general sequence

$$\bar{T}_{i+1} = \frac{T_i}{n} \sum_{\alpha=1}^N \sum_{j=1}^{n_\alpha} \frac{2\pi}{\phi_{j\alpha}(T_i)}, \quad (9)$$

where the actual time interval used is increased by the standard deviation  $\sigma_{i+1}$  of the mean,  $T_{i+1} = \bar{T}_{i+1} + \sigma_{i+1}$ . Adding one standard deviation guarantees that the majority of the local-

ized electrons have made at least one full revolution.

The coarse graining of time through the time intervals  $T_i$ , whose length is defined by the localized electrons, plays a crucial role for the description of the entire cluster dynamics. It provides the natural time scale to interpolate between the explicit time-dependent dynamics of the classical electrons and the time-averaged rate description of the bound “quantum” electrons. Over an interval  $t_0 < t < t_0 + T_i$  in time all processes involving quantum rates will be considered within a fixed cluster environment with properties averaged over the previous interval  $[t_0 - T_{i-1}, t_0)$ .

### III. COARSE-GRAINED PHOTOIONIZATION INTO THE PLASMA

We are now prepared to calculate atomic properties in the environment of other cluster ions and electrons with the understanding that all these processes are for a specific time interval  $T_i$  as introduced in the previous section. The photoionization dipole matrix elements for many-electron atoms provided within the Hartree-Fock approximation [17] allow one to determine the cross section for ionization of individual occupied orbitals to the continuum; see Sec. II A.

To apply these cross sections, we have to approximately map the present situation of a cluster ion surrounded by the localized electrons and other charged particles (ions and delocalized electrons) into an effective single-ion scenario. This requires one first to determine the electronic energy of an ion with its localized electrons and then to construct an energy-equivalent configuration.

#### A. Electronic energy of the ions including localized electrons

Averaged over  $T_i$  we calculate the number  $n_\alpha$  of electrons localized about ion  $\alpha$  and their mean energy

$$E_\alpha^* = E(q_\alpha) + \sum_{j=1}^{n_\alpha} \left( \frac{p_j^2}{2} - q_\alpha v(j, \alpha) \right) + \sum_{j>k=1}^{n_\alpha} v(j, k), \quad (10)$$

with  $v(j, k)$  the interaction potential, Eq. (8), between two particles of unit charge at positions  $\mathbf{r}_j$  and  $\mathbf{r}_k$ ,  $q_\alpha$  the charge of the ion  $\alpha$ , and  $E(q_\alpha)$  the energy of its bound electrons.

The localized electrons are in excited states of the ion  $\alpha$ , as shown in Fig. 3. Starting from the energy  $E^*$  of this actual configuration, we include them in the photoionization process by constructing the equivalent configuration of ion  $\alpha$ . In this configuration, we relax all localized electrons but one onto the last occupied orbital of the actual configuration. We put the remaining electron on a Rydberg orbit, whose energy is given by the condition that the actual and equivalent configurations have the same energy  $E^*$ .

#### B. Ionization potential of the equivalent configuration

To find the binding energy of electrons in occupied orbitals in the presence of a Rydberg electron, we assume that neither the quantum number  $n$  nor the angular momentum  $\ell$  of the Rydberg electron changes upon release of an electron from a deeper orbital. Then the ionization energy  $\Delta E$  is given by

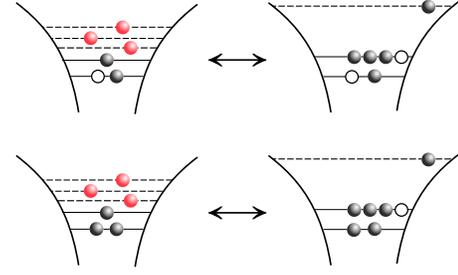


FIG. 3. (Color online) Sketch for the construction of the equivalent electronic configuration (right panel) from the actual configuration (left panel) of a cluster ion with three localized electrons (red) around a  $3s^1 3p^1$  (upper row) and  $3s^2 p^1$  (lower row) configuration. Holes are shown as open circles.

$$\Delta E(C_q, n) = E(C_{q+1}, n) - E(C_q, n), \quad (11)$$

where  $E(C_q, n)$  is the energy of the valence shell configuration  $C$  with charge  $q$  and an additional electron in a Rydberg orbital  $n$ . We have omitted  $\ell$  as an index in Eq. (11) since the following discussion does not depend on it. As an illustrative example, we will use  $\ell=1$ . The quantum number  $n$  can take values from  $n_0$  up to  $\infty$ , with  $n_0$  corresponding to the situation where the “Rydberg” electron is in the lowest possible state; for argon and  $\ell=1$  it is  $n_0=3$ . Figure 4 provides energies for six configurations and  $n=3, \dots, 9$ . In five of the configurations the valence shell does not contain electron holes. Thus they are solely defined by the charge  $C_q \equiv q$ . The case  $C'_2=3s^2 3p^4 np$ , also shown in Fig. 4, is an example for an exception with a hole in  $3s$ . Obviously, we have

$$E(C_q, \infty) = E(C_{q+1}, n_0) = E(C_q, n_0) + E_{\text{ip}}(C_q); \quad (12)$$

i.e., the asymptotic energy  $n \rightarrow \infty$  of a Rydberg series coincides on the one hand with the origin of the Rydberg series of

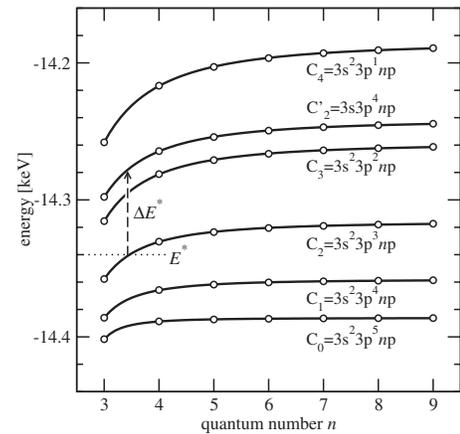


FIG. 4. Total energies (circles) for ions with one of the electrons in a Rydberg state with quantum number  $n$  and angular momentum  $\ell=1$ . The valence shell configuration  $C_q$  is specified for each set of energies. Fitted curves (thick lines) according to Eq. (13). As an example, the energy  $E^*$  for an equivalent configuration (dotted line) as defined in Sec. III A and the corresponding energy  $\Delta E^*$  for ionization (dashed line) are shown.

the next higher charge state, as can be seen in Fig. 4. On the other hand, it is equal to the sum of the ground-state energy and the ionization potential  $E_{\text{ip}}$  for an ion with configuration  $C_q$ . For finite values of  $n$  we approximate the energies by a quantum defect formula [22]

$$E(C_q, n) = E(C_q, \infty) - \frac{1}{2} \left( \frac{Z_{\text{eff}}}{n - \mu_q} \right)^2, \quad (13)$$

where in contrast to the usual ionic charge  $Z$  we use an effective one

$$Z_{\text{eff}} = (n_0 - \mu_q) [2E(C_q, \infty) - 2E(C_q, n_0)]^{1/2}, \quad (14)$$

chosen such that the first level ( $n_0$ ) of the series agrees with the exact value while the quantum defect  $\mu_q$  is fitted. Hence, Eq. (13) is very accurate at intermediate  $n$ , where we need it. When fitting the curves, Eq. (13), to the calculated energies (cf. Fig. 4), we found the quantum defects  $\mu_q$  to be almost independent of  $q$ . This allows us in the calculation of energy differences, Eq. (11), to eliminate the term containing  $n$  and  $\mu$  in Eq. (13). We get for a configuration  $C_q$  with an initial energy  $E^*$

$$\Delta E^*(C_q) = [E(C_{q+1}, \infty) - E^*] - b[E(C_q, \infty) - E^*], \quad (15)$$

with

$$b = \frac{E_{\text{ip}}(C_{q+1})}{E_{\text{ip}}(C_q)}, \quad (16)$$

the ratio of the ionization potentials; cf. Eq. (12). Thus we have obtained an expression for the energy necessary to ionize an electron from a valence shell in the presence of a Rydberg electron. This expression does not depend on the actual quantum numbers  $n$  and  $\ell$ , but only on the energy  $E^*$ .

### C. Condition for overbarrier inner ionization

Although we know now with Eq. (15) the ionization potential for a screened isolated ion, we have to position the ion in the cluster environment in order to decide if photoabsorption leads to photoexcitation within the ion  $\alpha$  or to inner ionization above the lowest barrier on the way to a neighboring ion. The energy balance for the photoelectron which decides between these two options is

$$E_i + V_i + \omega = E_f + V_f, \quad (17)$$

with local contributions from the ion to which the electron is bound (see Fig. 5),

$$E_i = -\Delta E^*, \quad E_f = E_{\text{kin}} - (q_\alpha + 1)v(\alpha, \text{bar}), \quad (18)$$

and contributions from the background of charges in the cluster,

$$V_i = -\sum_{j \neq \alpha} q_j v(j, \alpha), \quad V_f = -\sum_{j \neq \alpha} q_j v(j, \text{bar}), \quad (19)$$

where the index  $j$  runs over the delocalized electrons and all ions but ion  $\alpha$ . As introduced earlier,  $v(l, m)$  is the interaction between two Coulomb particles at positions  $\mathbf{r}_l$  and  $\mathbf{r}_m$ . Hence,  $V_i$  refers to the potential energy of the electron under

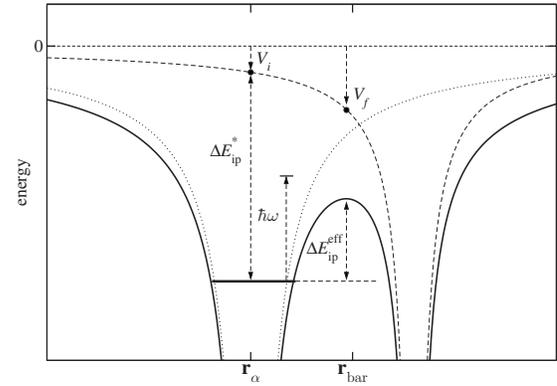


FIG. 5. Sketch of the energy balance for the photoionization of a cluster ion  $\alpha$ . The cluster environment is symbolized by a single neighboring ion. The thick solid line indicates the full cluster potential, and the thin long dashed line represents the cluster environment  $V(\mathbf{r})$ —i.e., the cluster potential without the contribution from the ion itself. The dotted line is the Coulomb potential for  $q_\alpha + 1$ —i.e., the field of the ion if the electron were ionized. The interaction of the bound electron with the nucleus is represented by  $\Delta E_{\text{ip}}^*$ . The saddle point of  $V(\mathbf{r}) - (q_\alpha + 1)v(\alpha, \mathbf{r})$  defines the position  $\mathbf{r}_{\text{bar}}$  of the barrier.

consideration and located at the position of its mother ion  $\alpha$  due to interaction with particles of charge  $q_j$  at  $\mathbf{r}_j$ . Likewise,  $V_f$  is the potential energy of the same electron at the potential barrier  $\mathbf{r}_{\text{bar}}$  due to the interaction with the same charged particles as before. The energy balance, Eq. (17), is taken with respect to the location  $\mathbf{r}_\alpha$  of the ion and the location  $\mathbf{r}_{\text{bar}}$  of the lowest potential barrier near the ion.  $\Delta E^{\text{eff}}$  in Fig. 5 is defined by putting  $E_{\text{kin}} = 0$  in Eq. (18)—i.e.,

$$\Delta E^{\text{eff}} = \Delta E^* - (q_\alpha + 1)v(\alpha, \text{bar}) + V_f - V_i. \quad (20)$$

Figures 6 and 7 give an overview of the coarse-grained variables during the laser pulse for an atom at the center and one at the surface of the  $\text{Ar}_{147}$  cluster. Figure 6 shows first of all an overview of their evolution during the whole interaction with the laser. Starting at the ground state of the neutral Ar, each absorbed photon leads to a rising step in the total energy. Note that for an ion in the cluster the electron has to be excited only above the lowest barrier. Moreover, the energy  $E^*$  of the equivalent configuration takes merely the localized electrons into consideration and not the newly ionized one. Therefore, each ionization event leads to jumps of  $E^*$  higher than the energy of the photon; cf. Fig. 6.

The electron localization is equivalent to a relaxation of the system and lowers therefore  $E^*$ . The flat regions observable for both ions correspond to the case where there are no localized electrons, when the total energy of the ion is given solely by the bound, “quantum” electrons. The smaller final charge of the surface atom is a consequence of the cluster expansion. The surface expands much faster than the core, leading to higher interionic barriers and an early suppression of the inner-ionization. A detailed view of the evolution of the two atoms is shown in Fig. 7 for the time interval from  $t = 143$  to 151 fs. The coarse graining is symbolized by the

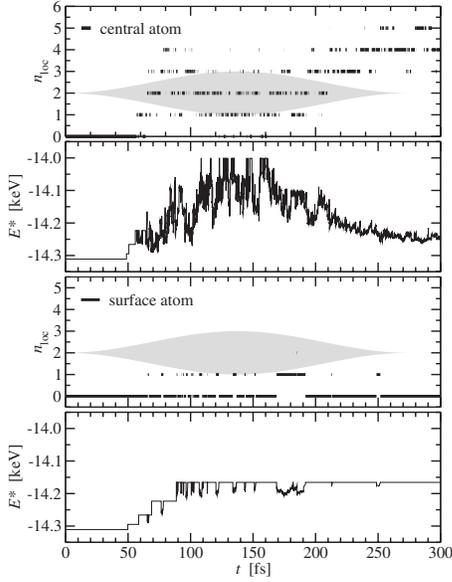


FIG. 6. Number of localized quasifree electrons  $n_{loc}$  and the corresponding energy  $E^*$  according to Eq. (10) for the central ion and a surface ion of an  $Ar_{147}$  cluster from the microscopic calculation in Sec. IV.

solid lines, showing the total energy  $E_\alpha^*$  of the two atoms averaged over the time intervals  $T_i$ , as described in Sec. II C.

#### D. Effective cross section for inner photoionization

Finally, we are in the position to adopt the photoionization cross section as formulated in Sec. II A for an isolated ion to the situation of an ion in the cluster. Here, we take the lowest potential barrier to a neighboring ion as effective ionization threshold. Therefore, the actual cross section as shown in Fig. 1 can vary between minimum and maximum possible

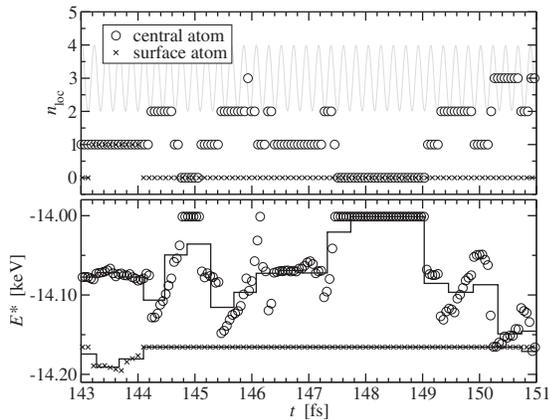


FIG. 7. Close-up of Fig. 6 for  $143 < t < 151$  fs. The upper panel shows the number of localized electrons ( $\circ$  for the central atom, where  $q=6$ , and  $\times$  for the surface atom,  $q=4$ ). The lower panel shows the total energy  $E^*$  of the ions and their localized electrons. The solid lines show here the average of  $E^*$  over the time intervals  $T_i$ , as introduced in Sec. II C. The gray line shows the electric field of the laser; it corresponds to the gray-shaded area of Fig. 6.

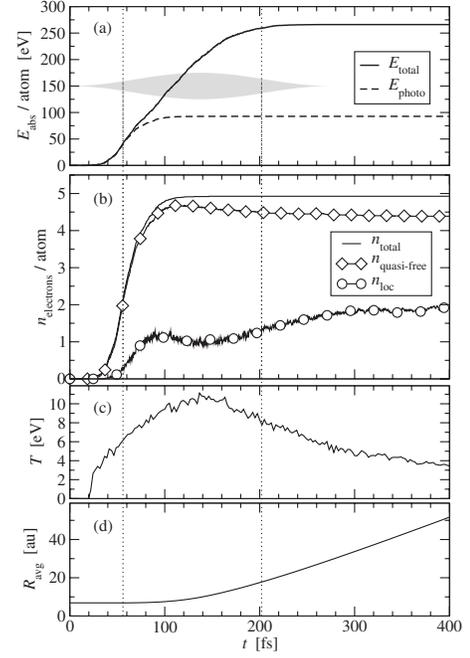


FIG. 8. Explosion of  $Ar_{147}$  exposed to a 100-fs-long vuv laser pulse with intensity  $I=7 \times 10^{13}$  W/cm<sup>2</sup> and photon energy  $\hbar\omega = 20$  eV: (a) solid line, total absorbed energy per atom; dashed line, energy absorbed due to photoionization; (b) solid line, total number of ionized electrons per atom;  $\diamond$ , quasifree electrons;  $\circ$ , localized electrons; (c) temperature of the quasifree electrons; (d) expansion, average interionic distance [see Eq. (22)].

values of potential barriers according to Eq. (5). The interval is indicated for small photon energies in Fig. 1 with the additional dotted line.

The electrons available for photoionization are only the tightly bound ones from the actual configuration, while the matrix elements entering the expression for the cross section, Eq. (3), take into account the screening of the electrons as provided by the equivalent configuration  $\sigma = \sigma(q_{eqv})$ . Hence, the multiplicity has to be taken from the actual configuration  $w_{act}$  to arrive at the screened photoionization cross section

$$\sigma^{scr}(q) = \frac{w_{act}}{w_{eqv}} \sigma(q_{eqv}). \quad (21)$$

#### IV. DYNAMICS OF $Ar_{147}$ UNDER AN INTENSE VUV LASER PULSE

We will illustrate the theoretical framework introduced above with the dynamics of  $Ar_{147}$  exposed to a 100-fs vuv laser pulse with  $\hbar\omega=20$  eV (i.e., wavelength  $\lambda=62$  nm) and an intensity of  $7 \times 10^{13}$  W/cm<sup>2</sup>. Figure 8 shows the main quantities characterizing the response of the cluster to the vuv pulse—namely, the energy absorption, the ionization degree, the temperature of the plasma, and the cluster explosion. The latter is characterized by the interionic distance  $R_{avg}$ , defined as

$$R_{\text{avg}} = \sqrt{\frac{1}{N} \sum_{\alpha=1}^N \min_{\beta \neq \alpha} (|\mathbf{r}_{\beta} - \mathbf{r}_{\alpha}|^2)}. \quad (22)$$

In the first part of the pulse, until approximately  $t=80$  fs, the absorption is dominated by photoionization. After that, inverse bremsstrahlung (IBS) sets in and photoionization starts to saturate. We can easily disentangle the two contributions to the total absorbed energy since we treat photoionization via rates (see Sec. II A) and IBS through classical propagation (see Sec. II B). The energy  $E_{\text{total}}$  shown in Fig. 8(a) is the sum of all photoionization events and the energy absorbed during the laser driven dynamics divided by the number of cluster atoms.

Note that until this time, as can be seen in panel (b) of Fig. 8, almost all electrons ( $n_{\text{total}}$ ) are trapped ( $n_{\text{quasifree}}$ ) inside the cluster due to their low kinetic energy. They start to leave the cluster with the onset of the IBS heating. Even when the photon frequency is larger than the first atomic ionization potential, not all atoms can be ionized as if they were separated. Rather, the space charge built up even in a relatively small cluster such as  $\text{Ar}_{147}$  allows the electrons to become *quasifree* only. Hence, the “ionization into the cluster”, as described in Sec. III, starts very early in the pulse. These electrons thermalize very quickly and obey a Maxwell-Boltzmann velocity distribution with a temperature  $T$ , which is shown as a function of time in Fig. 8(c).

Figure 8(b) also shows the average number  $n_{\text{loc}}$  of the quasifree electrons which are localized, as introduced in Sec. II C. They reach a maximum ( $t \approx 90$  fs) when the photoionization becomes unlikely. The increase of the temperature of the electron plasma [see Fig. 8(c)] from this point on favors a decrease of the number of localized electrons. At the point ( $t \approx 120$  fs) where the cluster expansion [see Fig. 8(d)] starts, the temperature of the electrons plasma decreases, despite the continuing energy absorption, and the localization increases again to an average of two electrons per atom.

This transient complex behavior of electrons and ions in the cluster may be accessible experimentally soon using attosecond pulse probe techniques; for a proposal, see [23]. For the time being, we present in Fig. 9 more conventional observables measurable in the experiment, such as the final charge distribution of ions and the kinetic-energy distribution of electrons for the dynamics of  $\text{Ar}_{147}$ . They have been obtained by propagation up to  $t=6.4$  ps—i.e., much longer than shown in Fig. 8. This time is long enough to ensure that the ions are so far separated that electron exchange between them is negligible. Those electrons which are (classically) recombined would stay at the corresponding ion if the propagation would be extended. The dominating fragment is  $\text{Ar}^{3+}$  despite the fact that about five electrons [Fig. 8(b)] per atom have been photoionized. The kinetic energy distribution of the electrons can be fitted with an exponential decay  $\exp(-E_{\text{kin}}/E_0)$  with  $E_0=5.4$  eV, thus emphasizing the ther-

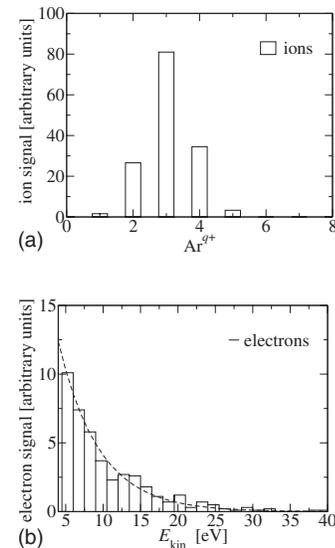


FIG. 9. Ion charge distribution (top panel) and kinetic energy distribution of the released electrons (bottom) after the cluster explosion for the same pulse as in Fig. 8. The propagation time was  $t=6.4$  ps. The electron signal can be exponentially fitted by  $\exp(-E_{\text{kin}}/E_0)$  with  $E_0=5.4$  eV (dashed line).

mal origin of the electrons. The temperature of the plasma has similar values; see Fig. 8(c).

In comparison with an earlier approach [4], which completely neglects the effect of the screening electrons, one observes as a general trend that the number of quasifree electrons is considerably smaller. This is connected with a weaker absorption of energy. E.g., for the example discussed above, only about five electrons per atom are (inner) ionized as compared to eight using the earlier approach [4] and the total absorbed energy is less by about 50%.

## V. SUMMARY

We have described a quantum-classical hybrid approach to follow the time-dependent dynamics of rare-gas clusters exposed to intense short laser pulses. Special attention has been paid to incorporate the screening of cluster ions by the electron plasma, formed by quasifree electrons which have been ionized from their mother ions, but cannot leave the cluster as a whole due to the strong background charge. The mean time scale of these localized quasifree electrons is determined and provides the link between the microscopic dynamics of all charged particles and the quantum dynamics of photoionization which is described by ionization rates adopted to the screening in the cluster environment.

Hence, this approach is especially well suited to tackle interaction of clusters with light from vuv and x-ray free-electron laser sources. As an illustrative example, we have discussed the dynamics of  $\text{Ar}_{147}$  exposed to a 100-fs laser pulse of  $I=7 \times 10^{13}$  W/cm<sup>2</sup> intensity and  $\hbar\omega=20$  eV photon energy.

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