

Data assimilation by means of synchronization

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Dresden, 29 - 31 July 2009

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The model and
synchronization
scheme

Surface growth
picture and
mean-variance
diagram

Perfect model limit:
Noiseless case

Non perfect model:
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problem

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This contribution is being published in J. Geophys. Res.,
doi:10.1029/2009JD012411,
<http://www.agu.org/journals/pip/jd/2009JD012411-pip.pdf>

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- Yang et. al. (2006) J. Atmos. Sci.: “Lorenz 3-d model”
- Duane et. al., (2006) Nonlin. Proc. Geophys.: “Lorenz 3-d model, geostrophic model”
- Cohen et. al. (2008) Phys. Rev. Lett.: “High-dimensional chaos optical experiment”

Assimilation by means of synchronization: the truth

$$\dot{\mathbf{x}}_T = f(\mathbf{x}_T) + \xi_{\text{model}}$$

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Assimilation by means of synchronization:
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and the model

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$$\dot{\mathbf{x}}_M = f(\mathbf{x}_M) + \kappa(\mathbf{x}_T + \xi_{\text{obv}} - \mathbf{x}_M)$$

where ξ_{obv} , ξ_{model} represent noisy observations and
model errors, respectively

Our “reality”: Lorenz '96 model

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This toy model mimics some aspects of the dynamics of the atmosphere such as advection, constant forcing, and linear damping.

$$\partial_t u_n = -u_{n-1}(u_{n-2} - u_{n+1}) - u_n + F + \eta_n(t)$$

$n = 1 \dots L$, $L = 256$ is the system size and $F = 8$

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 $\eta_n(t)$ is a delta-correlated Gaussian noise representing the fast degrees of freedom that our model of reality will not be able to describe

$$\langle \eta_n(t_1) \eta_m(t_2) \rangle = 2D \delta_{nm} \delta(t_1 - t_2)$$

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Euler algorithm with time step $\Delta t = 10^{-4}$ and periodic boundary conditions.

Our computer (forecasting) model

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- ▶ Two steps: 1) Assimilation (synchronization) for a given time. 2) Prediction, \mathbf{v} evolves freely.

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- ▶ Two steps: 1) Assimilation (synchronization) for a given time. 2) Prediction, \mathbf{v} evolves freely.
- ▶ Prediction time horizon $\sim \lambda_{\max}^{-1}$.

Synchronization error and the surface picture

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- ▶ We want to analyze the space-time correlations of $\mathbf{w}(t)$

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- ▶ We want to analyze the space-time correlations of $\mathbf{w}(t)$

- ▶ We define the “surface”

$$h_n(t) = \ln |w_n(t)| = \ln |u_n(t) - v_n(t)|$$

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$$h_n(t) = \ln |w_n(t)| = \ln |u_n(t) - v_n(t)|$$
- ▶ The spatial structure of the surface h_n : Power spectral density (PSD) or structure factor
$$\mathcal{S}(q, t) = \langle \hat{h}_q(t) \hat{h}_{-q}(t) \rangle$$

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 $h_n(t) = \ln |w_n(t)| = \ln |u_n(t) - v_n(t)|$
- ▶ The spatial structure of the surface h_n : Power spectral density (PSD) or structure factor
 $S(q, t) = \langle \hat{h}_q(t) \hat{h}_{-q}(t) \rangle$
- ▶ For a generic perturbation the surface exhibits scale-invariant correlations below the correlation length $\ell_c(t) \sim 1/q_c(t)$

$$S(q, t) \sim \begin{cases} q^{-(2\alpha+1)} & \text{if } q \gg q_c(t) \\ a(t) & \text{if } q \ll q_c(t) \end{cases} . \quad (1)$$

Rough surfaces

- ▶ We say that h_n is a rough surface with a *roughness exponent* α and a correlation length $l_c(t)$

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- ▶ Errors in chaotic (dissipative and homogeneous) systems scale as Kardar-Parisi-Zhang (KPZ) surfaces: $\alpha = 1/2$ in $d = 1$.
- ▶ For dissipative chaos: $\partial_t w(x, t) = \partial_{xx} w + \xi(x, t)w$
- ▶ $h \equiv \ln |w|$ leads to

$$\partial_t h(x, t) = \partial_{xx} h + (\partial_x h)^2 + \xi(x, t)$$

The KPZ equation for surface growth.

Mapping errors to rough surfaces

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- ▶ Many important quantities can be mapped into well-known magnitudes in terms of the surface

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Mapping errors to rough surfaces

- ▶ Many important quantities can be mapped into well-known magnitudes in terms of the surface

- ▶ The (squared) *surface width*

$W^2(t) = \left\langle (1/L) \sum_{n=1}^L [h_n(t) - \bar{h}]^2 \right\rangle$, informs about the the correlation length ℓ_c :

$$W^2(t) \propto \int S(q, t) dq \sim \ell_c(t)^{2\alpha}. \quad (2)$$

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$$W^2(t) \propto \int S(q, t) dq \sim \ell_c(t)^{2\alpha}. \quad (2)$$

- ▶ This allows us to obtain the **extent of the spatial correlations** from the simple measurement of the surface width!!

Mapping errors to rough surfaces

- ▶ The error *amplitude* or size:

$\varepsilon(t) = \langle \|\mathbf{w}(t)\|_0 \rangle = \langle \exp \left[\bar{h}(t) \right] \rangle$, informs about the
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Mapping errors to rough surfaces

- ▶ The error *amplitude* or size:

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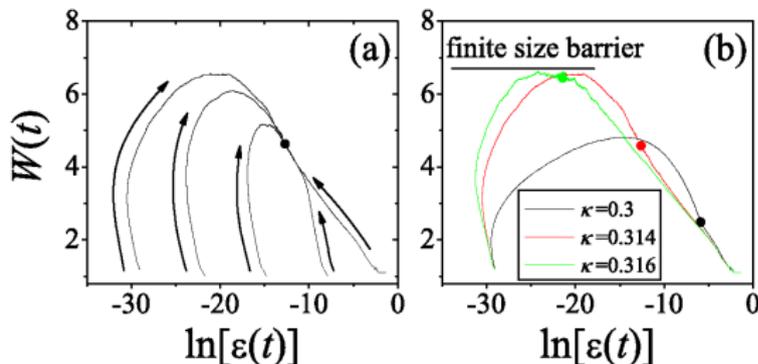
- ▶ We have $\varepsilon(t) \sim \exp(\lambda_{\max} t)$, where λ_{\max} is just the average surface velocity ($\bar{h} = \lambda_{\max} t$) and also corresponds to the first Lyapunov exponent.

The mean-variance diagram

- ▶ Phase diagram of perturbations: surface width vs. average height \iff spatial correlation length (localization magnitude) vs. perturbation size

The mean-variance diagram

- ▶ Phase diagram of perturbations: surface width vs. average height \iff spatial correlation length (localization magnitude) vs. perturbation size
- ▶ The mean-variance of the logarithm of perturbations (MVL) diagram: $W^2(t)$ vs. $\ln \varepsilon(t)$



Perfect model limit: Noiseless case

- ▶ The model describes reality “perfectly”

Reality:

$$\partial_t u_n = -u_{n-1}(u_{n-2} - u_{n+1}) - u_n + F + \cancel{\eta_n(t)}$$

Model:

$$\partial_t v_n = -v_{n-1}(v_{n-2} - v_{n+1}) - v_n + F + \kappa(u_n - v_n)$$

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- ▶ No time scales are poorly described: Noise intensity $D = 0$
- ▶ Perfect synchronization ($\|w\| \rightarrow 0$) for large enough coupling, $\kappa > \kappa_c$,
- ▶ Partial synchronization ($\|w\|$ finite) below the threshold, $\kappa < \kappa_c$, with a typical length scale ξ_x

Spatial correlations of errors: Power spectral density

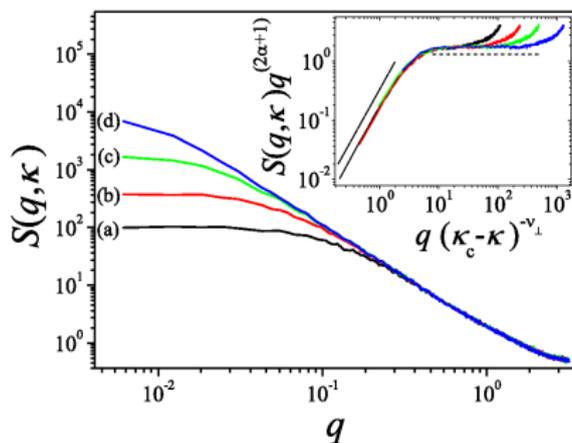
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- ▶ Error surface $h = \ln |w|$
- ▶ roughness exponent is KPZ: $\alpha = 1/2$
- ▶ $q_x \sim \xi_x^{-1} \sim |\kappa - \kappa_c|^\nu$,
as $\kappa \rightarrow \kappa_c^-$.
- ▶ the exponent $\nu \approx 0.85$
is also universal.



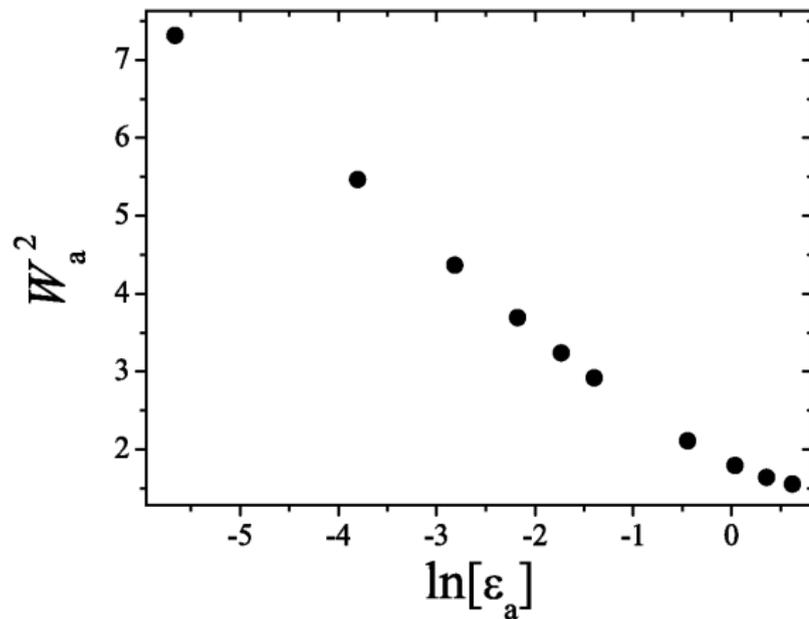
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MVL diagram: Fixed point



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The model only describes reality partially:

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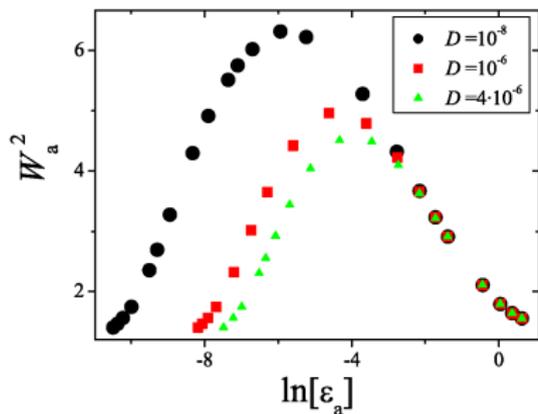
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Asymptotic fixed point for $D \neq 0$



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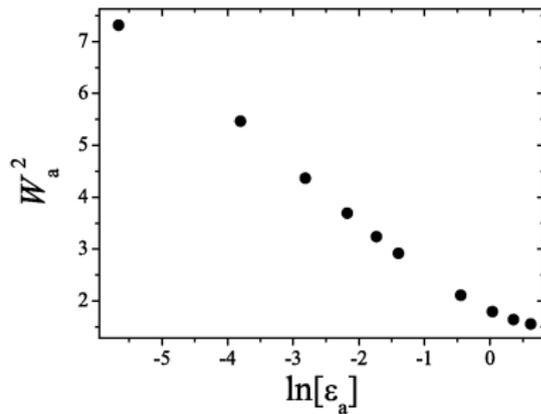
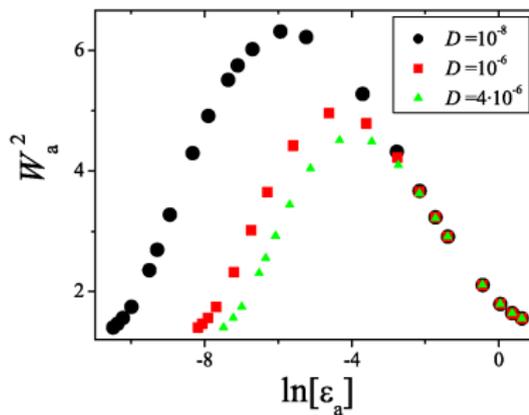
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$D = 0$

Spatial correlations of model-reality errors

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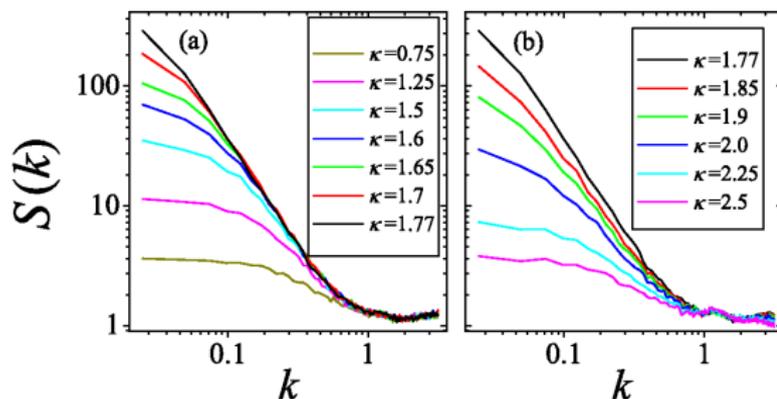
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$\kappa < \kappa_{\text{opt}}$

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Optimal coupling

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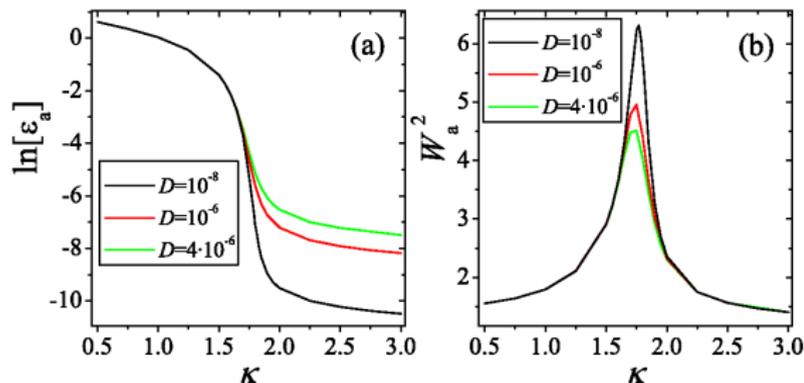
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Free evolution of assimilated trajectories

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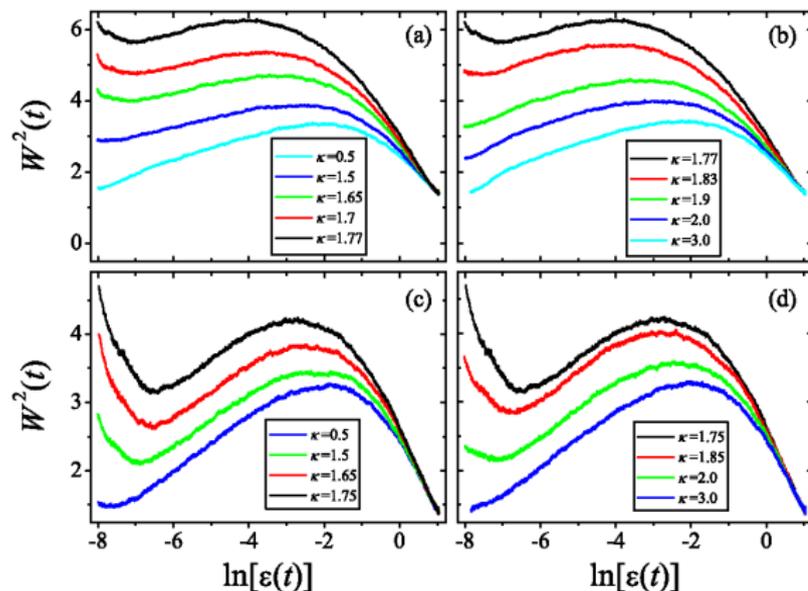
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a,b: $D = 10^{-8}$ and c,d: $D = 10^{-6}$

Workshop announcement

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Exploring Complex Dynamics in High-Dimensional Chaotic Systems: From Weather Forecasting to Oceanic Flows

MPIPKS Dresden, 25 - 29 January 2010

Scientific Coordinators: Juan M. López, Arkady Pikovsky, and Antonio Politi

Organisation: Claudia Pöenisch (MPIPKS Dresden, Germany)

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